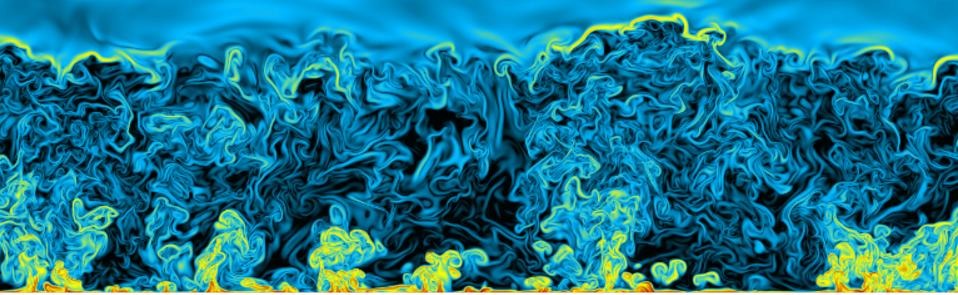


#### **BOUNDARY LAYER METEOROLOGY**



#### **Prof. Ivana Stiperski, Dr. Manuela Lehner** Department of Atmospheric and Cryospheric Sciences

### **Chapter 3**

#### Statistical treatment of turbulence



#### Content

3.1. Averaging, Stationarity and Homogeneity

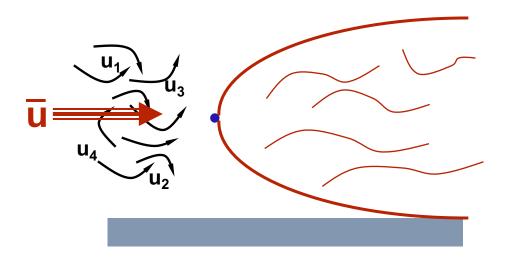
- 3.2. Taylor Hypothesis
- 3.3. Reynolds Decomposition
- 3.4. Co-variances and their Physical Meaning
- 3.5. Other Turbulence variables



# Why do we use statistics when dealing with turbulence?



#### **Motivation**



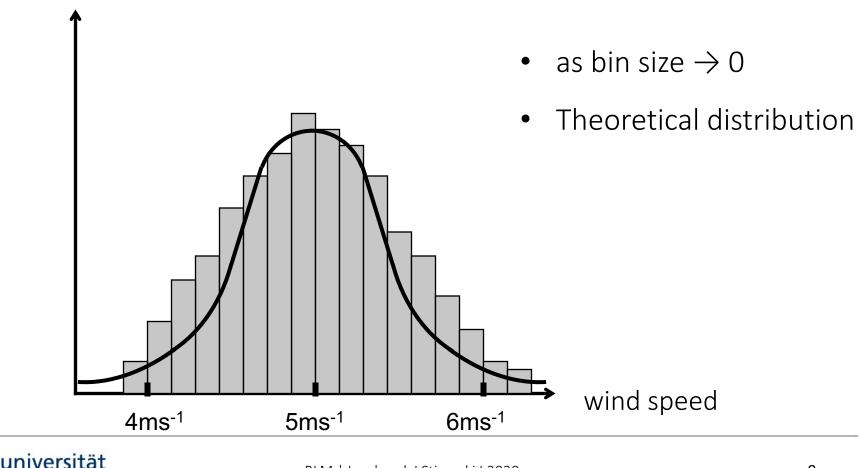
- irregular, different each second...
- cannot / do not want: describe each trajectory
- statistical treatment!



### Distributions

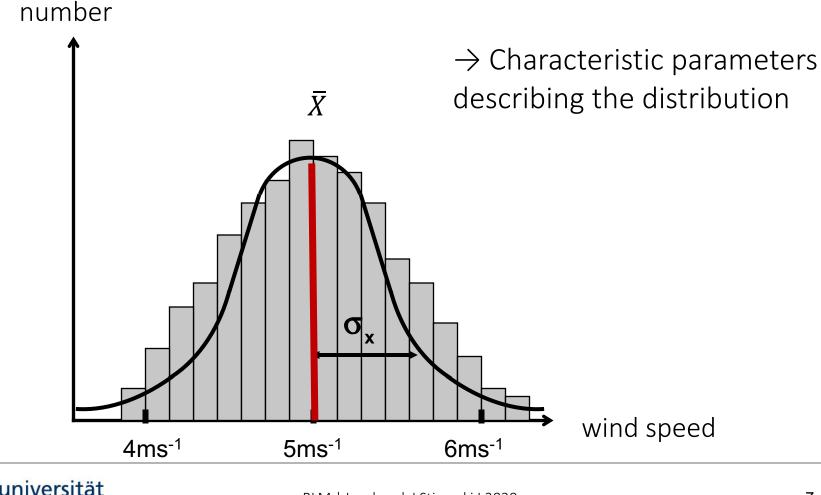
Represent measurements depending on their likelihood

number

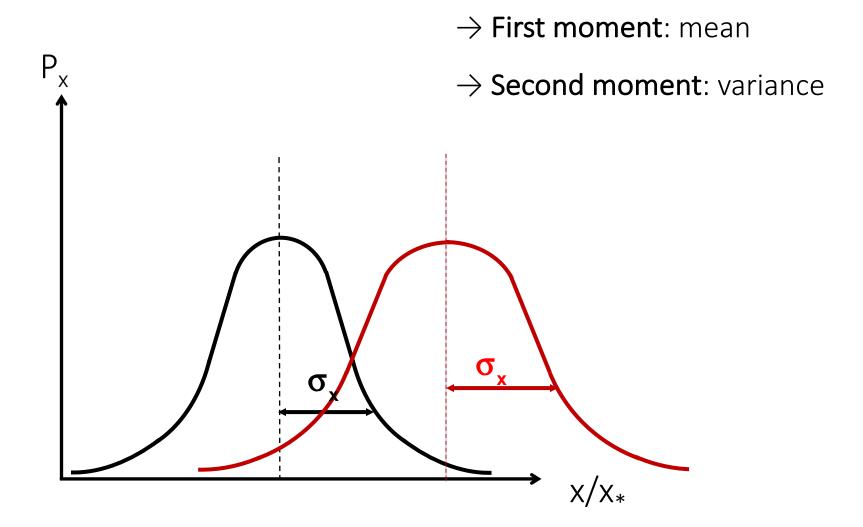


### Distributions

 $\rightarrow$  Likelihood presented through: Probability Density Function(PDF, P)



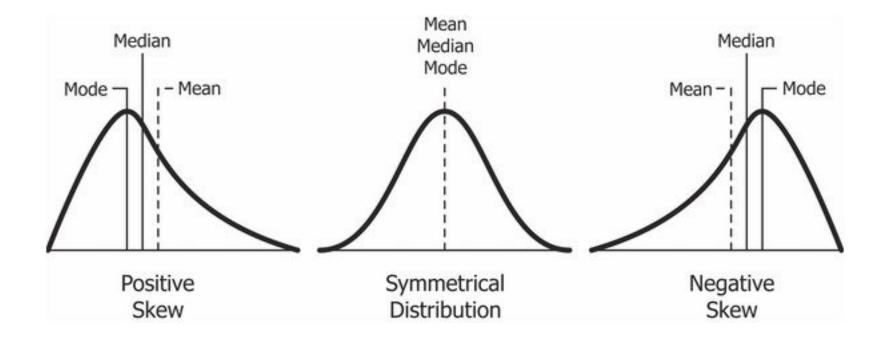
#### **Distributions - Variance**





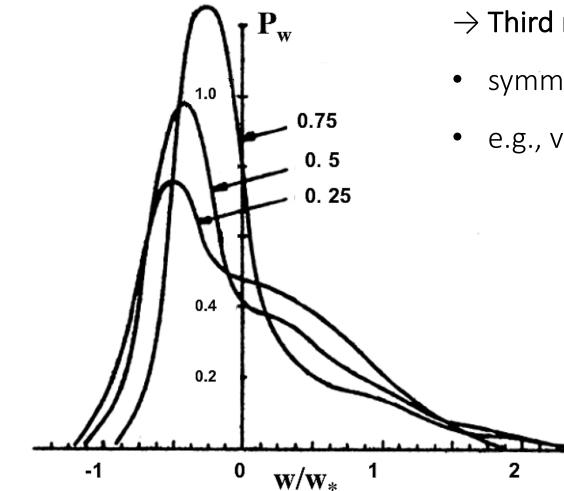
#### **Distributions - Skewness**

#### $\rightarrow$ Third moment: Skewness





### **Distributions - Skewness**

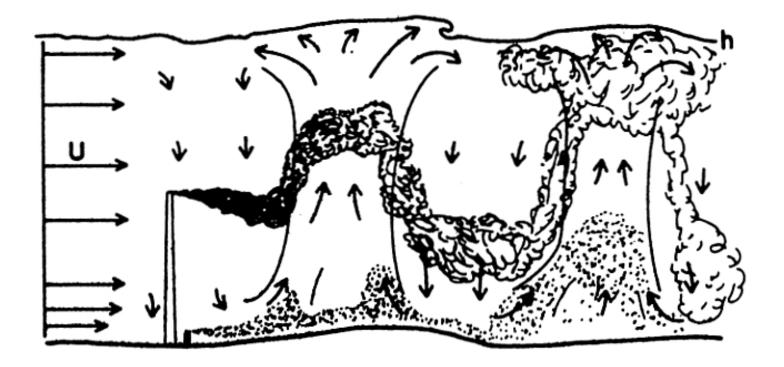


#### → Third moment: Skewness

- symmetric?
- e.g., vertical velocity in CBL



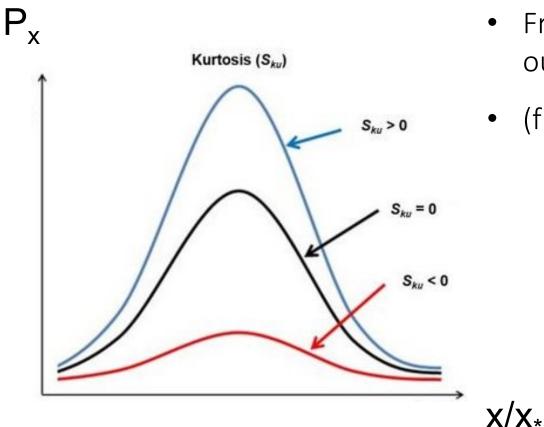
#### CBL – skewness w



- $\rightarrow$  single 'thermals' with strong positive w
- ightarrow occupy about 30-40% of surface
- $\rightarrow$  median (w) slightly negative



### **Distributions - Kurtosis**



#### ightarrow Fourth moment: Kurtosis

- Frequency of extremes outliers
- (flatness of a distribution)



### **Probability density function**

- completely characterized by its moments
- average:
- generally:

$$\overline{a} = \int_{-\infty}^{\infty} aP_a da$$
$$\overline{f(a)} = \int_{-\infty}^{\infty} f(a)P_a da$$

 $\infty$ 

- moments:

$$\overline{a^{n}} = \int_{-\infty}^{\infty} a^{n} P_{a} da$$
$$\overline{(a - \overline{a})^{n}} = \int_{-\infty}^{\infty} (a - \overline{a})^{n} P_{a} dx$$

 $-\infty$ 

• central moments:  $\rightarrow$  n=0: norm  $\rightarrow$  n=1:=0  $\rightarrow$  n=2: variance  $\rightarrow$  n=3: skewness

### **Normal Distribution**

$$P_a = \frac{1}{\sqrt{2\pi\sigma_a}} \exp\left\{-\frac{1}{2} \frac{(a-\overline{a})^2}{\sigma_a^2}\right\}$$

Moments of the normal distribution

- $\rightarrow$  skewness = 0
- $\rightarrow$  kurtosis = 3
- $\rightarrow$  all higher moments = 0



#### Tools



#### 1. Stationarity

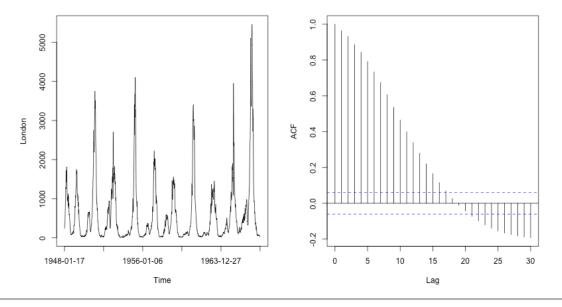


#### **Stationarity**

#### Consider: time correlation of a variable with itself

$$a(t) \cdot a(t') = C_a(\tau, T)$$

$$\tau$$
 = t - t'  
T= abs. time





#### **Stationary turbulence**

stationary turbulence 
$$\leftrightarrow a(t) \cdot a(t') =: C_a(\tau)$$

- $\rightarrow$  C<sub>a</sub> independent of T
- $\rightarrow\,$  for all  $\tau$
- $\rightarrow$  in particular also for  $\tau = 0$
- $\Rightarrow a(t) \cdot a(t) =: C_a(0) =$ variance



#### **Stationary turbulence**

 $\rightarrow$  variance independent of T

order of stationarity:

ightarrow correlation of the N-th order

$$a(t_1) \cdot a(t_2) \cdot a(t_3) \cdots a(t_N) = C_a^N(\tau_1, \tau_2, \tau_3, \dots, \tau_N)$$

- $\rightarrow$  independent of T
- ightarrow in particular: for au = 0

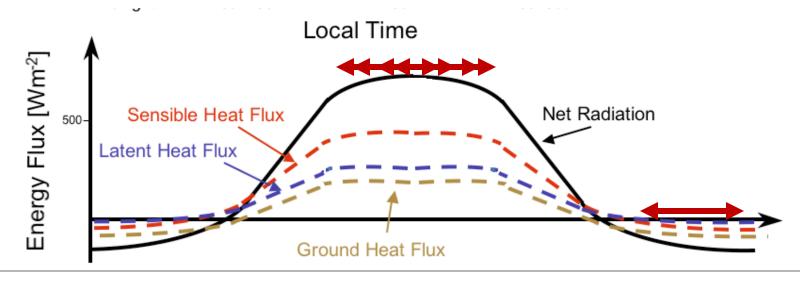
ALL moments of the distribution independent of T in practice:

average, variance independent of T 'enough'



### **Quasi - stationarity**

- turbulence is never really stationary
   → turbulence is *dissipative*
- 2 time scales
- $T_f$  = 'forcing time scale': external processes  $\rightarrow$  can be several hours



#### **Quasi - stationarity**

Turbulence is never really stationary
 --> Turbulence is dissipative

#### 2 time scales

- $T_f$  = 'forcing time scale': external processes
- $T_m$  = change of mean characteristics  $\rightarrow$  how long until 'mean profile' has adapted to (external) change

Quasi-stationarity:

$$T_m \ll T_f$$

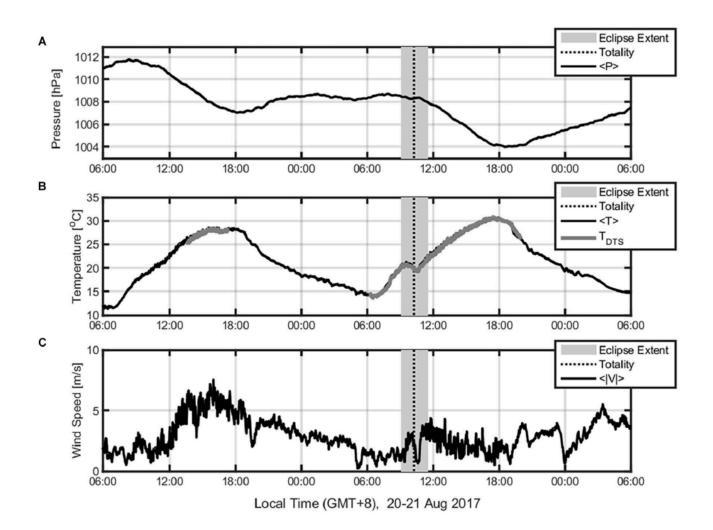


Higgins et al.2019: Ensemble-Averaging Resolves Rapid Atmospheric Response to the 2017 Total Solar Eclipse

https://www.frontiersin.org/articles/10.3389/feart.2019.00 198/full

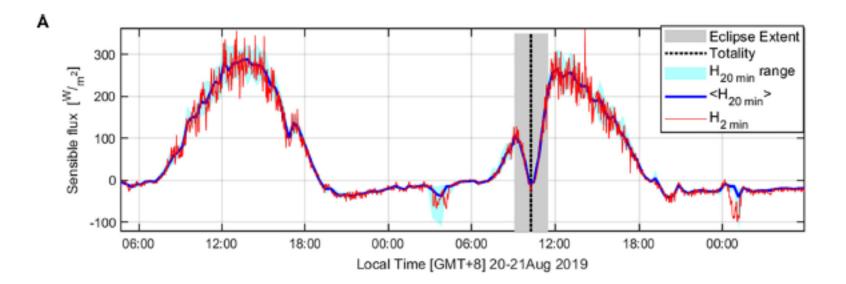


### Example: Solar eclipse



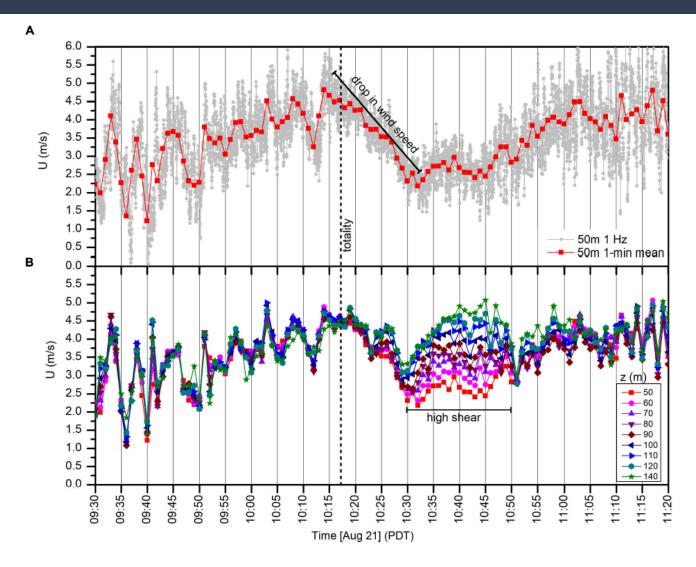


#### Example: Solar eclipse





#### Example: Solar eclipse





### Tools



#### 1. Stationarity

2. Homogeneity

ightarrow analogue of stationarity in space

$$a(x) \cdot a(x') =: C_a(\Delta x, \vec{r})$$
  
 $\vec{r} = \text{position vector}$ 

 $\rightarrow$  homogeneous:

$$a(x) \cdot a(x') = C_a(\Delta x)$$

ightarrow for all higher moments

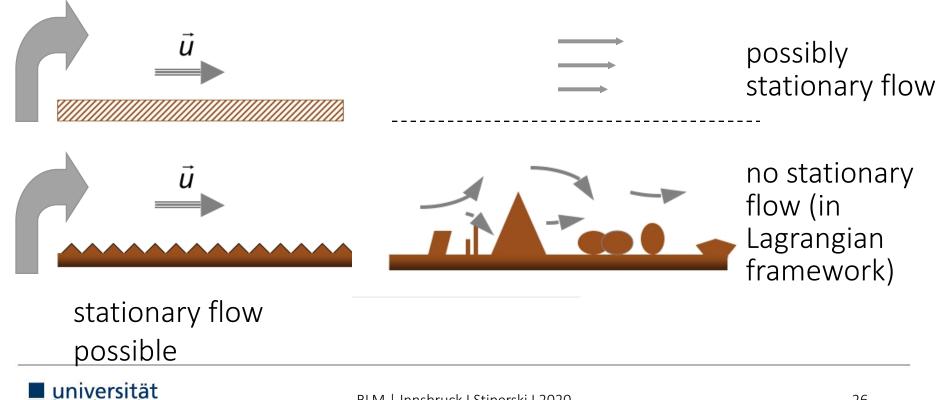
 $\rightarrow$  in particular  $\Delta$ x=0

independent of  $\vec{r}$ 



#### Homogeneity Stationarity

homogeneity corresponds to stationarity in space (horiz)  $\rightarrow$  if (sfc) forcing = const.



sbru

### Tools



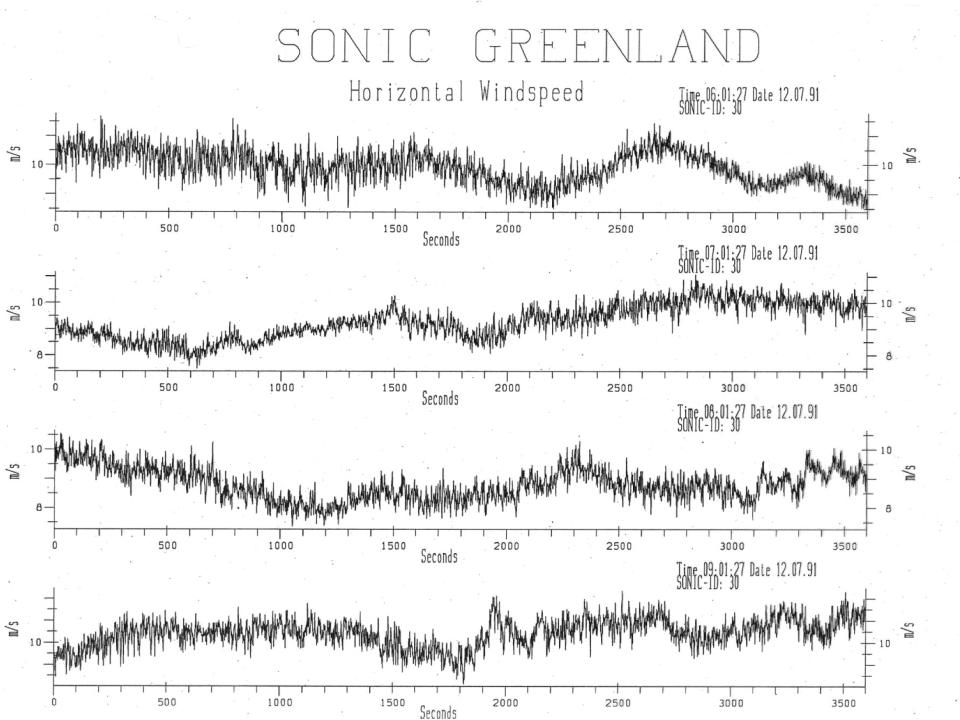
- 1. Stationarity
- 2. Homogeneity
- 3. Averaging



#### Averaging

• goal (e.g. of a measurement):





#### Averaging

- goal (e.g. of a measurement):
   → not: 'value' for this time at this place
  - → would like: to learn something about the physical processes, which produce such a time series
  - $\rightarrow$  to this time series, there is one 1m aside / '17 minutes' earlier, later, ...
  - → the random process (producing *this* time series) both spatially as well as in the time domain *theoretically goes to infinity*



#### Ensemble averages

## **Ensemble** = all possible realizations that can appear for a stationary process

 $\rightarrow$  average over all realizations:

$$\overline{a}^e = \frac{1}{N} \sum_{i=1}^N a_i(\vec{x},t) \text{ for } a = a(\vec{x},t)$$

atmosphere:  $N = \infty$ 



### Ergodic Hypothesis

E.H: under certain conditions:

$$\overline{a}^{x} \to \overline{a}^{e}$$
$$\overline{a}^{t} \to \overline{a}^{e}$$

Wyngaard (2010):

'The property that the time average of a stationary random variable and the space average of a homogeneous random variable converge to the ensemble average is called ergodicity'.

Ergodicity will always (implicitly) be assumed in real applications

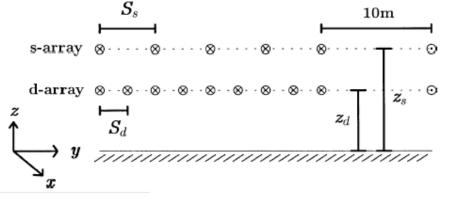


#### Averaging

Spatial average:

 $\overline{a}^{x} = \frac{1}{S} \iint_{S} a(\vec{x}, t) ds$  $\rightarrow$  want to know (structure of the turbulence)  $\rightarrow$  difficult to obtain (lots of instruments!)





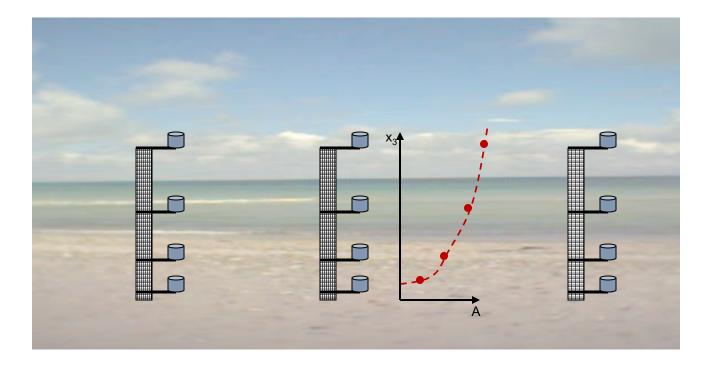
 $\rightarrow$  if horizontally homogeneous: one characteristic profile!

HATS field campaign  $\rightarrow$  Horst et al 2004  $\rightarrow$  small-scale turbulence



#### Averaging

#### → if horizontally homogeneous: one characteristic profile!





### Spatial Averaging

#### Fiber Optic Distributed Sensing (FODS)





#### Spatially Integrated Measurements





### Averaging

spatial average:

$$\overline{a}^x = \frac{1}{S} \iint_{S} a(\overline{x}, t) ds$$

time average:

$$\overline{a}^t = \frac{1}{T} \int_{t_1}^{t_1+T} a(\vec{x},t) dt$$

 $\rightarrow$  result of a measurement:

 $\rightarrow$  often: determine

often: determine 
$$\overline{a}^t \longrightarrow \overline{a}^e$$
 resp.  $\overline{a}^x \longrightarrow$  Ergodic Hypothesis!



### Averaging Rules

Table 3.1: Useful rules for averaging

A, B are variables, c is a constant

$$\overline{C} = C$$

$$\left(C \cdot A\right) = C \cdot \overline{A}$$

$$\left(\overline{A}\right) = \overline{A}$$

$$\left(\overline{\overline{A} \cdot B}\right) = \overline{A} \cdot \overline{B}$$

$$\left(\overline{A \cdot B}\right) \neq \overline{A} \cdot \overline{B}$$

 $\overline{\left(A+B\right)} = \overline{A} + \overline{B}$  $\overline{\left(\frac{\partial A}{\partial x}\right)} = \frac{\partial \overline{A}}{\partial x}$ 

An average behaves like a constant

The average of a product is not, in general, the product of the averages

This is an important property and derives from the Leibnitz theorem.



# Intermediate Summary



#### • Pdf's

- $\rightarrow$  probability density function to describe the variables
- $\rightarrow$  fully characterized through its *moments*
- stationarity
  - ightarrow all moments do not change with time
  - ightarrow in practice: up to second moments enough
- homogeneity
  - ightarrow is stationarity in space
- averaging
  - $\rightarrow$  would need: average over all possible realizations
  - $\rightarrow$  ensemble average
  - $\rightarrow$  certain conditions: time/space average  $\rightarrow$  ens. av.



# Tools

- 1. Stationarity
- 2. Homogeneity
- 3. Averaging
- 4. Taylor Hypothesis



• Mostly: have time series (one instrument, i.e. place)



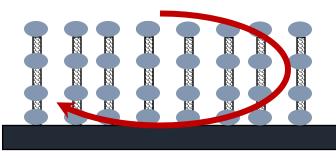


 Mostly: have time series (one instrument, i.e. place)
 → want information on the *structure of turbulence*

More generally:



 $\rightarrow$  how can I observe 'an eddy'?  $\rightarrow$  would need 1000's of instruments



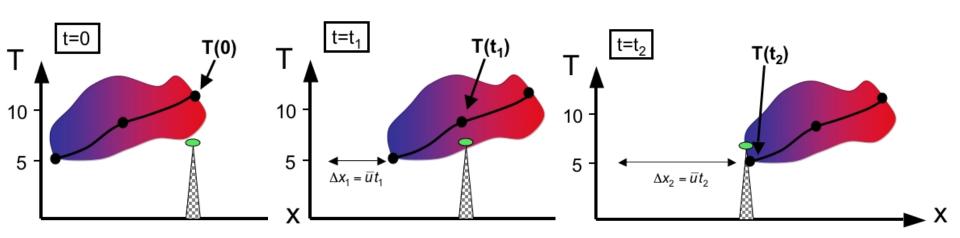


### Hypothesis:

The turbulence can be assumed to be **frozen** during the time it travels across the point of observation.

 $\rightarrow$  Taylor's Frozen Turbulence Hypothesis  $\rightarrow$  Geoffrey I. Taylor, 1938







### Hypothesis:

The turbulence can be assumed to be **frozen** during the time it travels across the point of observation.

→ applies, if:  
$$T_f >> L_e / \overline{u}$$
  
→ in practice:

$$\sigma_u / \overline{u} < 0.5$$

 $\rightarrow$  process is stationary

 $\rightarrow$  T<sub>f</sub> = forcing time scale

$$\rightarrow L_{e}$$
= characteristic length

$$\rightarrow$$
 u = average wind speed

 $ightarrow \sigma_{\scriptscriptstyle u}~$  = measure of activity of turbulence

$$\rightarrow \overline{u}$$
 = measure of advection



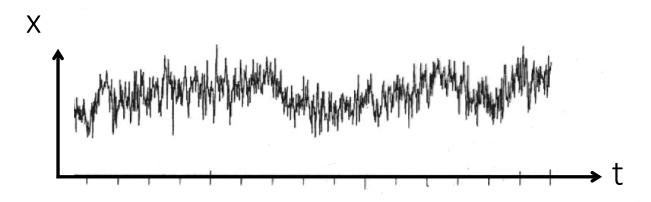
mathematically:  $D\zeta / Dt = 0$  $\frac{D\zeta}{Dt} = \frac{\partial\zeta}{\partial t} + \frac{\partial\zeta}{\partial x}\frac{\partial x}{\partial t} + \frac{\partial\zeta}{\partial y}\frac{\partial y}{\partial t} + \frac{\partial\zeta}{\partial z}\frac{\partial z}{\partial t} = 0$   $u \qquad v \qquad w \qquad (frozen')$   $\frac{\partial\zeta}{\partial t} = -\vec{v} \cdot \nabla\zeta \qquad (conservative flow field')$ 

→ measured change in time corresponds to advected spatial structure



# **Reynolds Decomposition**

averaging

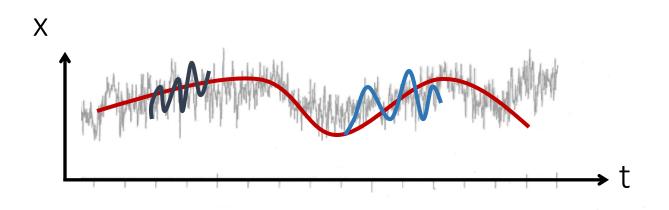


- $\rightarrow$  is it now turbulence?
- $\rightarrow$  or average flow?
- ightarrow averaging over what time?
- → where does turbulence remain after averaging?



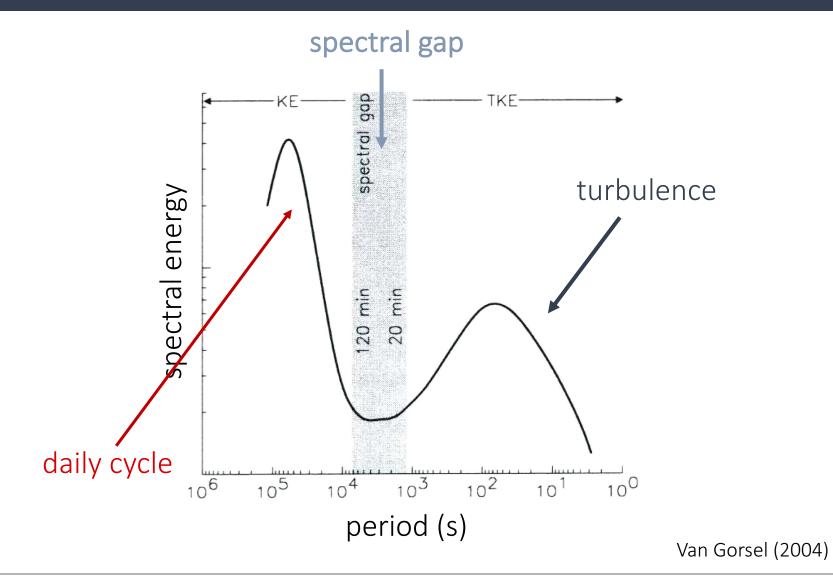
# **Reynolds Decomposition**

fluctuations with different periodicity
 → seasonal cycle, daily cycle, fast fluctuations



- spectral distribution
  - ightarrow how much 'power' in which periodicity?
  - $\rightarrow$  (see later, chapter 7)

# (Idealized) Energy Spectra

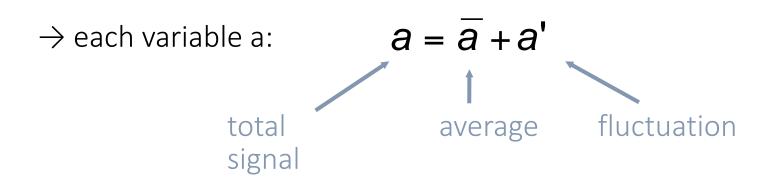




# **Reynolds Decomposition**

pragmatic choice:

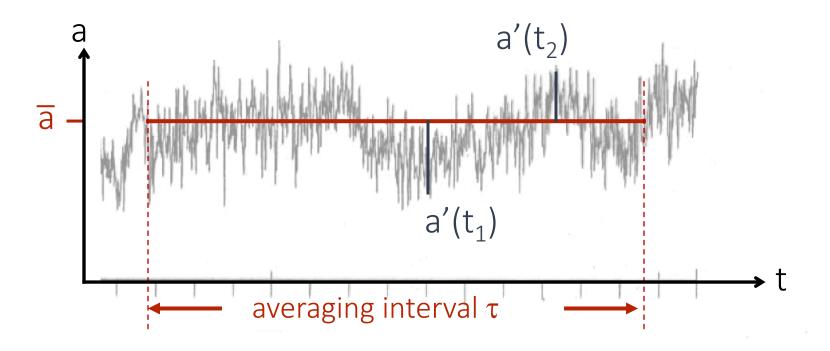
 → fluctuations faster than about 1h: turbulence
 → longer: average flow



*a* = time averageactually: ensemble average



# **Reynolds** Decomposition



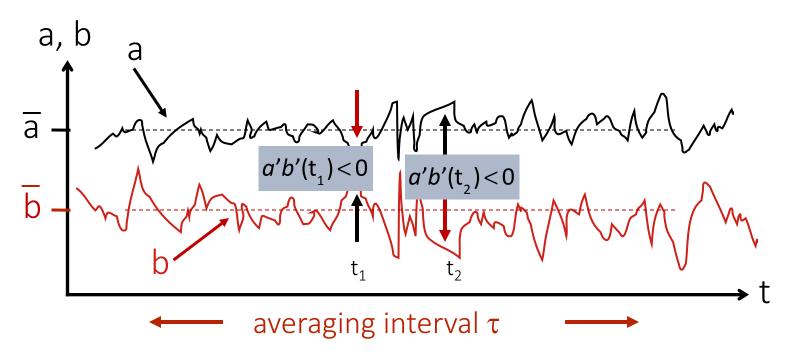
#### in practice:

- measure a(t): time resolution big enough (how to choose it?)
- compute: a
- for each averaging period
- from there: a'(t): for each measurement (20Hz)

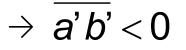
### Computation Rules for Reynolds Decomposition

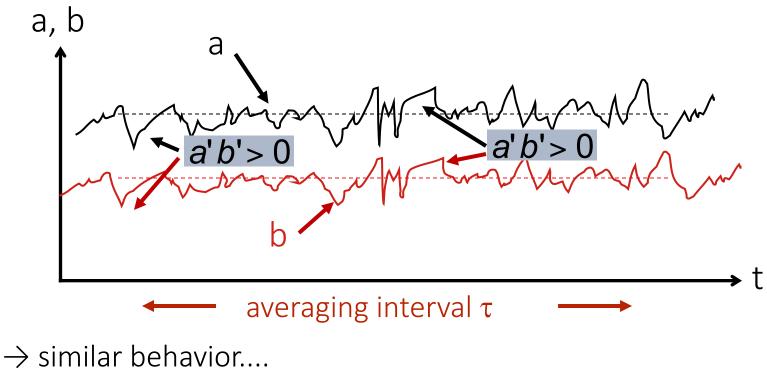
Table 3.2: Calcu	ulus for Reynolds Decomposition
<u>a</u> and b are variables, for which: $a = \overline{a} + a'$ ; $b = \overline{b} + b'$	
$\overline{a'} = 0$	By definition
$\overline{(a)} = \overline{(\overline{a} + a')} = \overline{a}$	By definition and 1)
$\overline{\left(\overline{b}\cdot a' ight)}=\overline{b}\cdot\overline{a}'=0$	The average of a product involving a primed variable vanishes
$\overline{(a \cdot b)} = \overline{(\overline{a} + a') \cdot (\overline{b} + b')}$ $= \overline{a} \cdot \overline{b} + \overline{a'b'}$	The covariance is not necessarily zero
$\overline{a^2} = \overline{a}^2 + \overline{a'^2}$	The second term on the <u>rhs</u> corresponds to the Second central moment, i.e. the variance
	and <i>b</i> are variable $\overline{a'} = 0$ $\overline{(a)} = \overline{(\overline{a} + a')} = \overline{a}$ $\overline{(\overline{b} \cdot a')} = \overline{b} \cdot \overline{a'} = 0$ $\overline{(a \cdot b)} = \overline{(\overline{a} + a')} \cdot \overline{(\overline{b} + b')}$ $= \overline{a} \cdot \overline{b} + \overline{a'b'}$





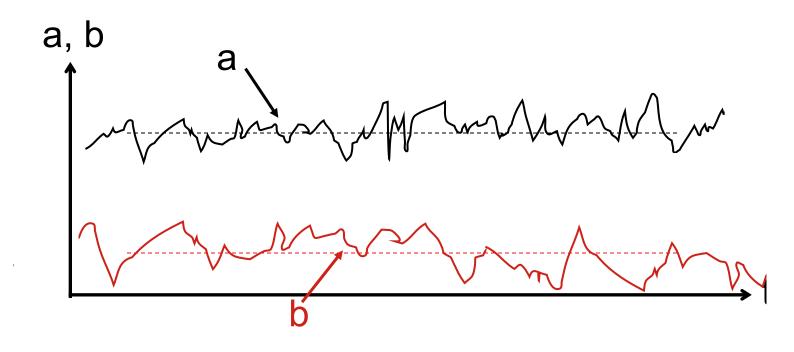
- $\rightarrow$  opposite behavior...
- $\rightarrow$  on **average** (here):





 $\rightarrow$  on average (here):

### $\rightarrow a'b' > 0$



- ightarrow a,b poorly correlated
- → a'b' ≈ 0



- physical meaning
   → turbulent transport
- in general:

 $\rightarrow$  consider: physical description of transport



### Description of Transport

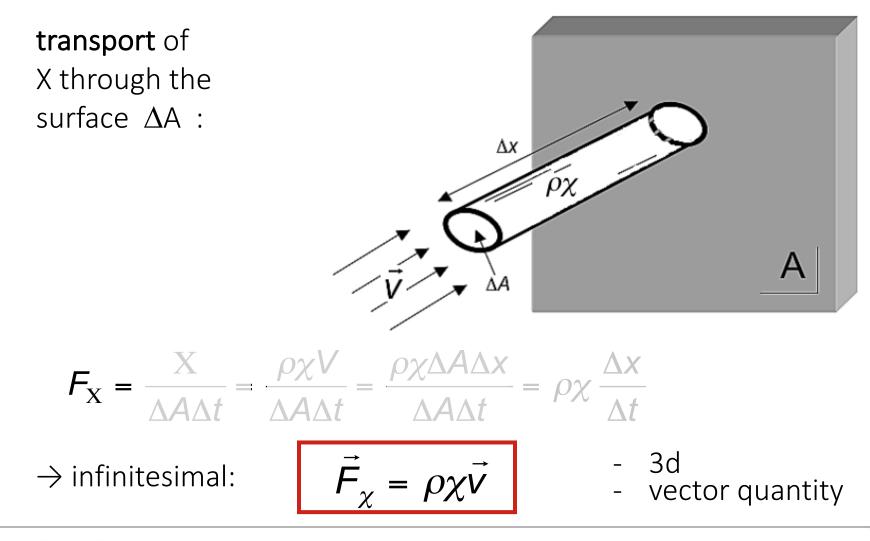
let: X = additive quantity ('countable')

then 
$$X = \iiint_V \rho \chi dV$$
 [X] = x

$$\chi$$
 = specific quantity  $[\chi] = x/kg$ 



### **Description of Transport**



# **Description of Transport**

 $\rightarrow$  infinitesimal:

$$\vec{F}_{\chi} = \rho \chi \vec{v}$$

 $\rightarrow$  Flux =  $\rho$  density of fluid

- x specific transported quantity
- v transport velocity
- $\rightarrow$  co-variances: **a'b'**  $\rightarrow$  often: a or b is a velocity component
- → for example: co-variance: a'w'→ transport of 'a' in vertical direction



## Covariances = Turbulent Transport

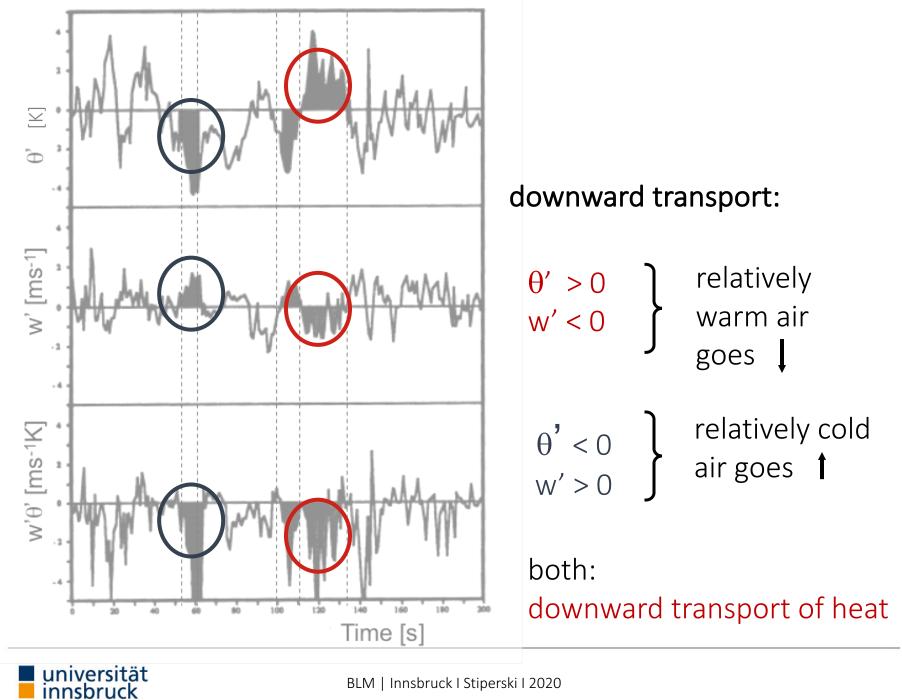
• these co-variances will be important:

 $w'\theta', w'q', u'w', (v'w')$ 

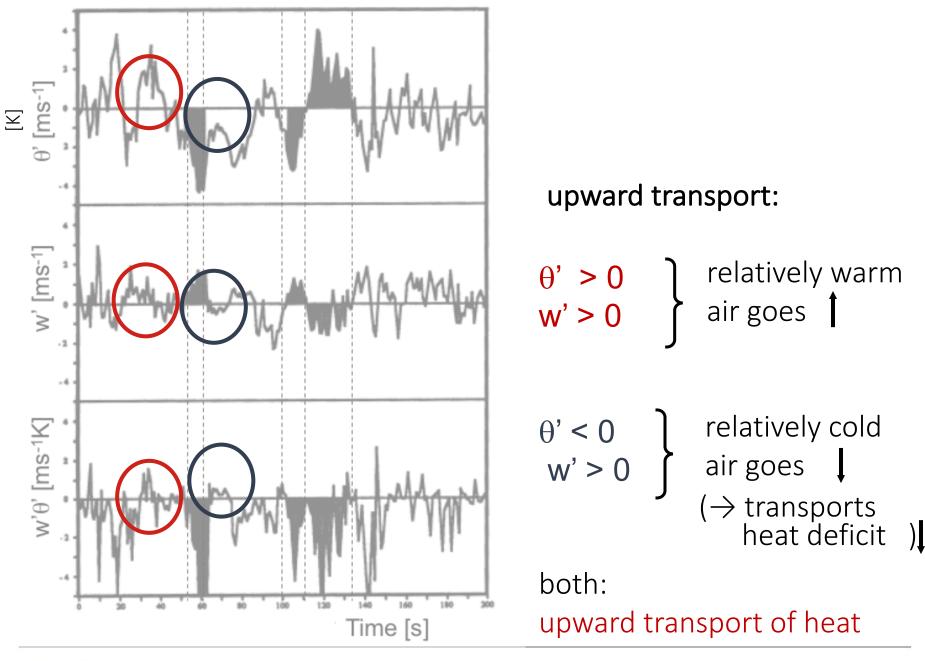
- with: w' and a scalar (θ', q'), resp. w' and u' → expresses transport of scalar quantities in vertical direction
- each instantaneous value, e.g. :
  - $\rightarrow$  instantaneous flux (transport)  $W'\theta'$

 $\rightarrow$  average: random or systematic, e.g.:

 $w'\theta'$ 



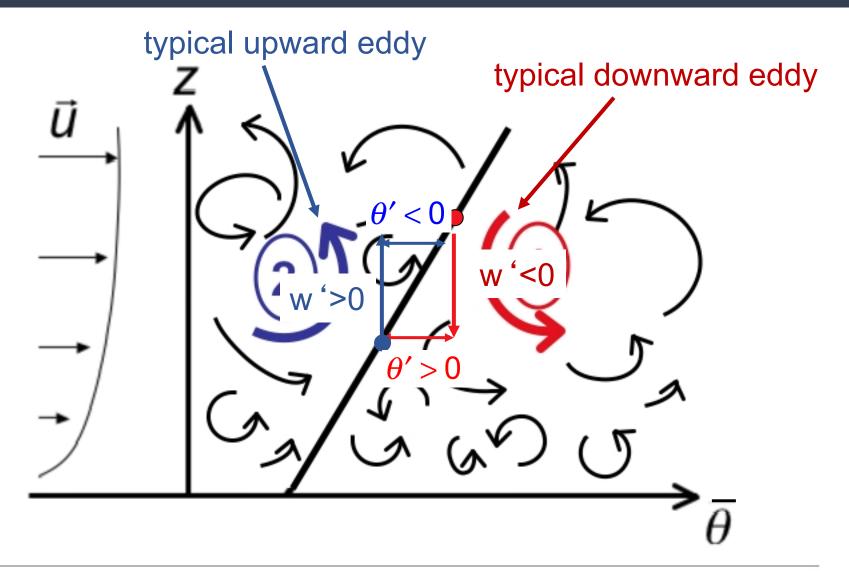
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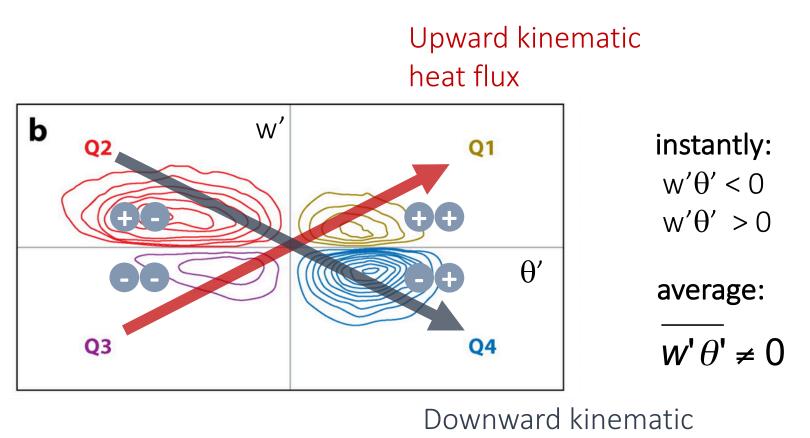
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## **Turbulent Fluxes**





### Quadrants

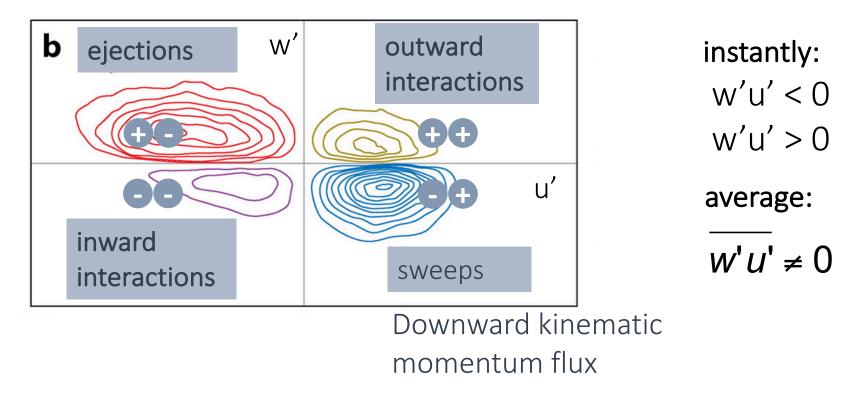


heat flux



### Quadrants

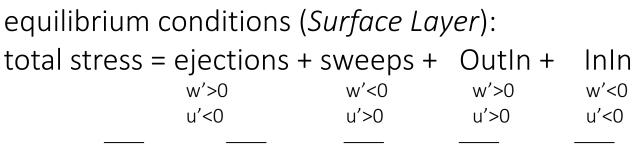
#### Upward kinematic momentum flux



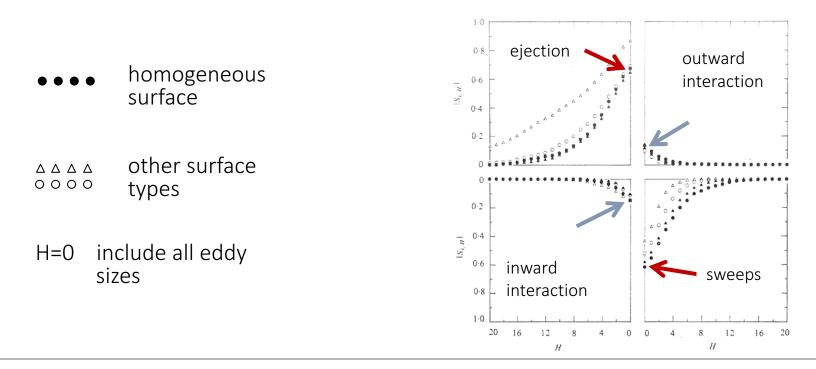
 $\rightarrow$  investigate type of transport



# Quadrant analysis: Momentum Transport

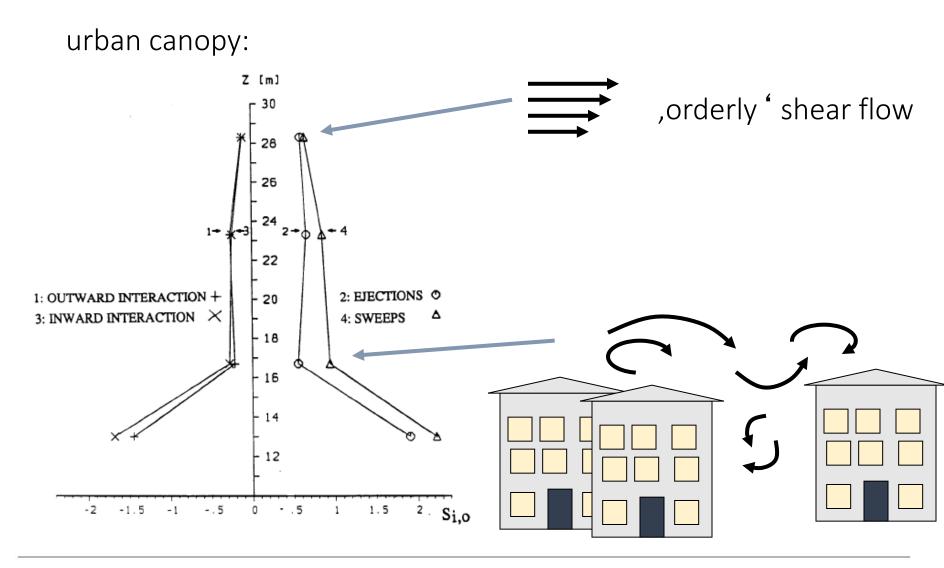


 $u'w'_{tot} = 0.6u'w'_{tot} + 0.6u'w'_{tot} - 0.1u'w'_{tot} - 0.1u'w'_{tot}$ 





# Quadrant analysis: Momentum Transport



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